



Radio Core Properties of Radio Quiet BAT AGN

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● 22 GHz 1" resolution survey of 100 radio-quiet BAT AGN with JVLA

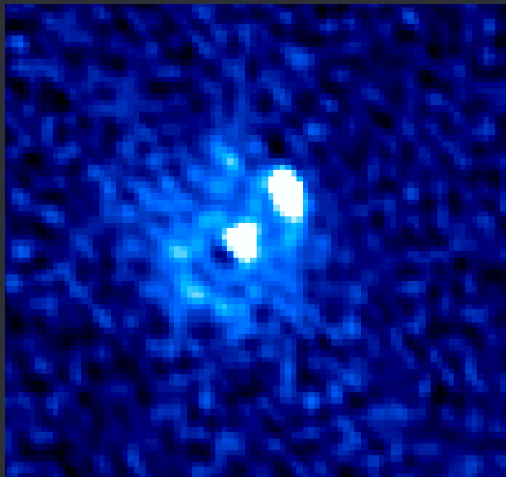
2015: Full BAT Sample → 313 *Herschel*-imaged → 48 radio-quiet AGN

2016: + 22 lower redshift BAT AGN

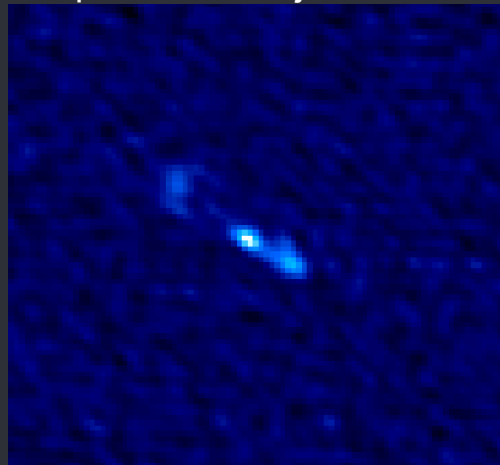
2017: +30 additional radio-quiet BAT AGN with low SFRs per unit stellar mass

Radio morphologies fall into 3 groups:

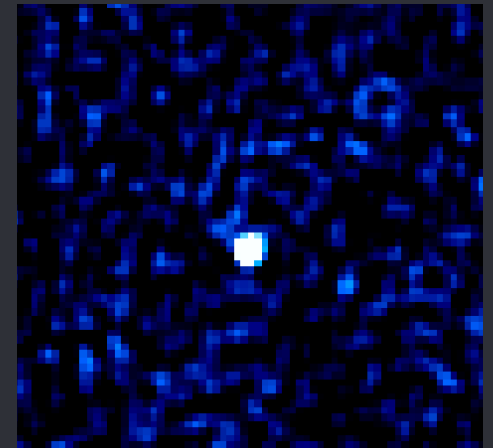
Extended Star Formation



Kiloparsec-scale jets

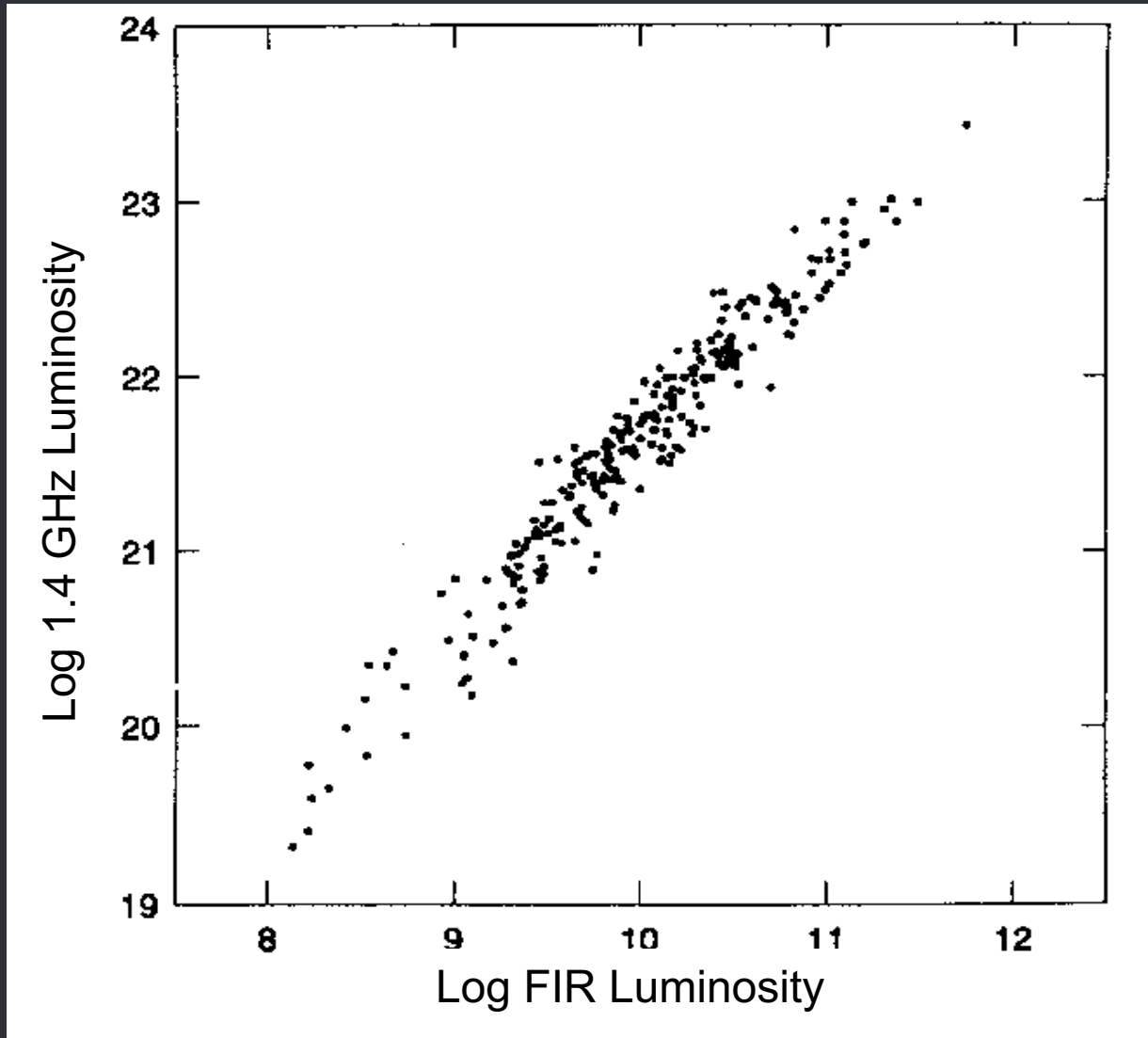


Unresolved / compact



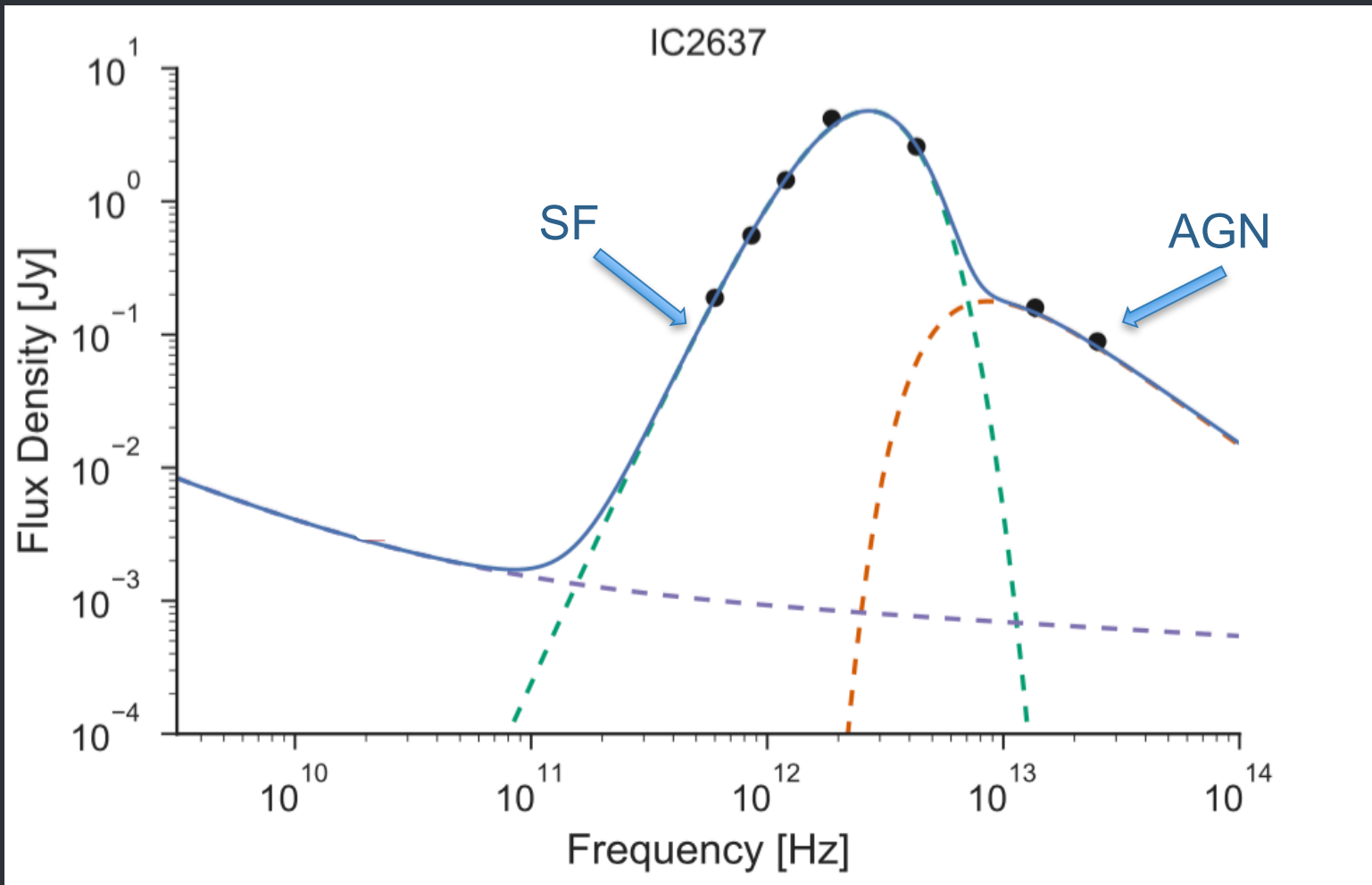
+ 4 non-detections

- Original Motivation: Do SFRs in AGN follow the FIR-Radio Correlation?



Condon 1992

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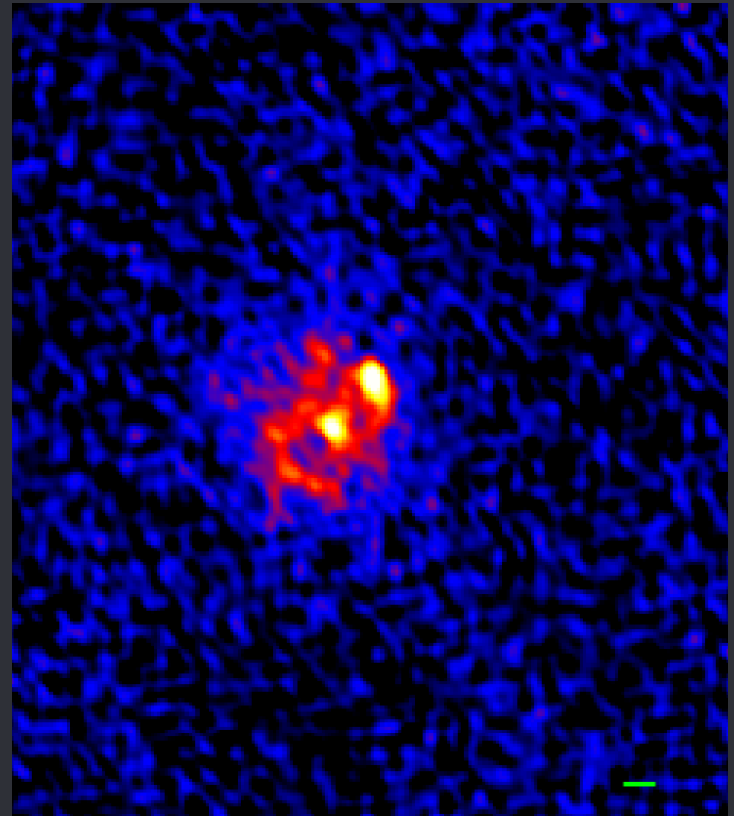


- HIGH RESOLUTION DECONVOLVES AGN FROM STAR FORMATION

Herschel (FIR)



VLA C-Array

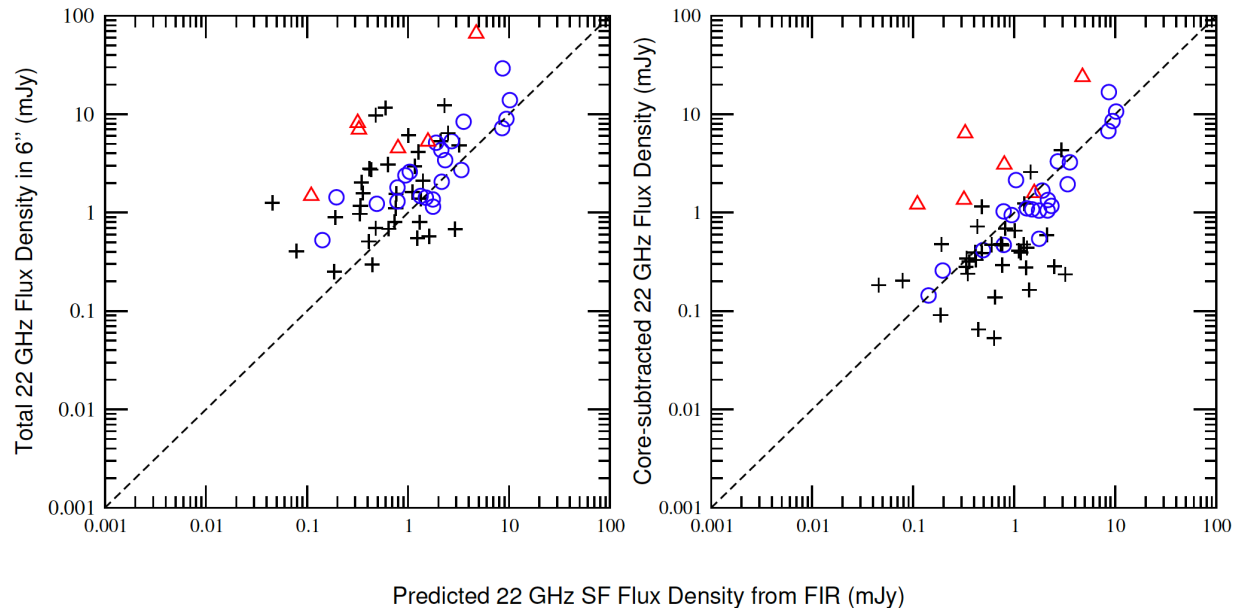


NGC 7679

SFRs in AGN *do* follow the FIR-radio correlation once the SF components are removed in *both* wavebands.

Total flux in image:

Extended flux only:



***But all those nice high-res images
also gave us an isolated radio core!***

***And the origin of radio emission in
radio-quiet AGN is an open debate:***

Dense star formation?

Tiny, scaled-down jets?

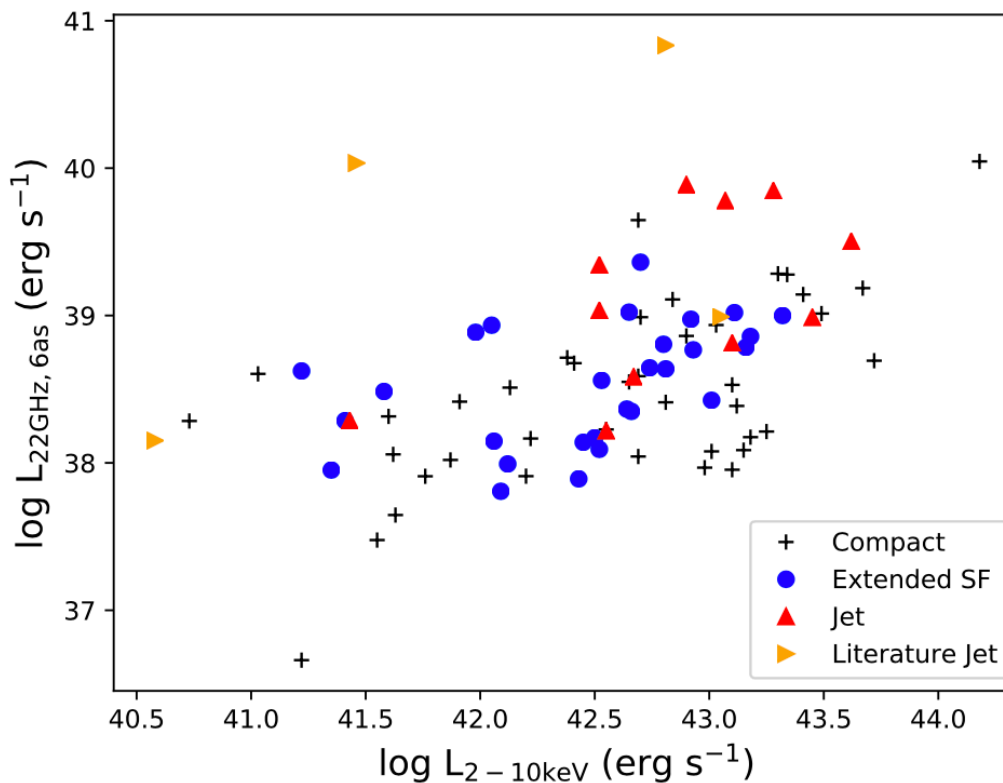
A stellar-like corona?

Sopp & Alexander 1991
Terlevich 1992

Heinz & Sunyaev 2003
Miller et al. 1993

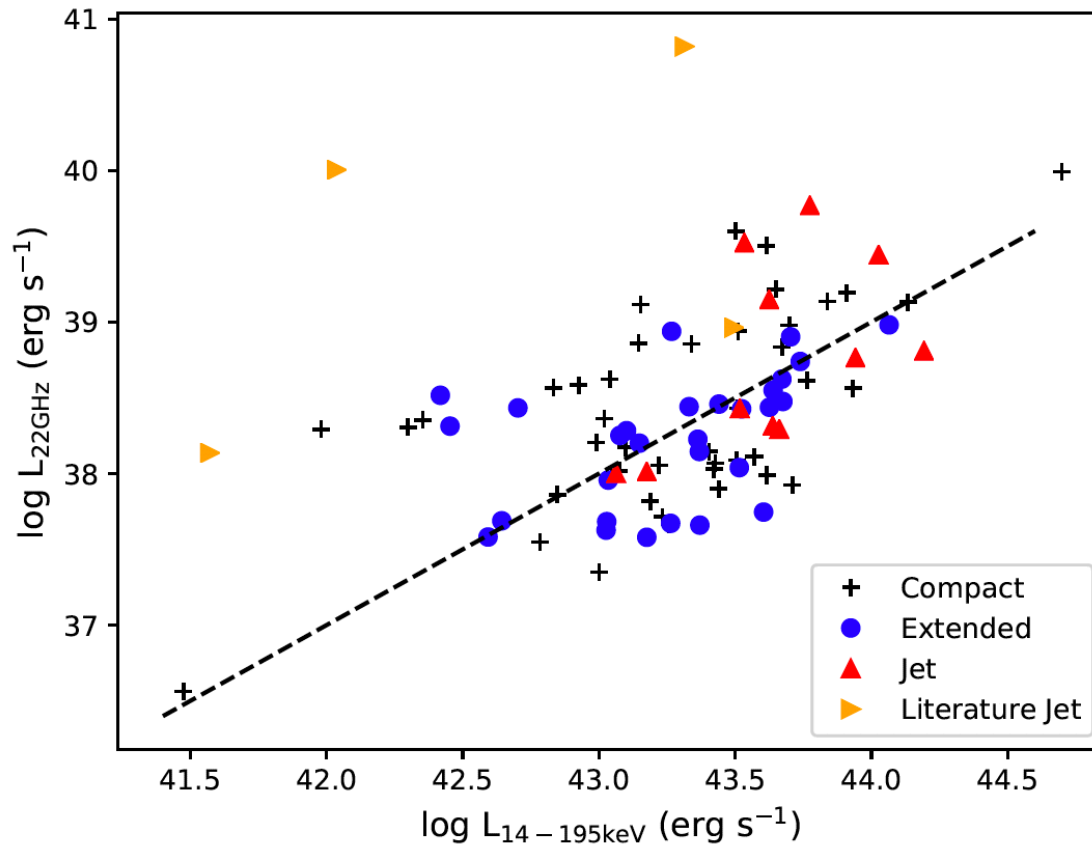
Laor & Behar 2008

People often use L_R / L_X as a diagnostic in low-resolution observations



There is *no diagonal line* above which you can confidently say that the radio emission within a 6" beam definitively comes from a jet.

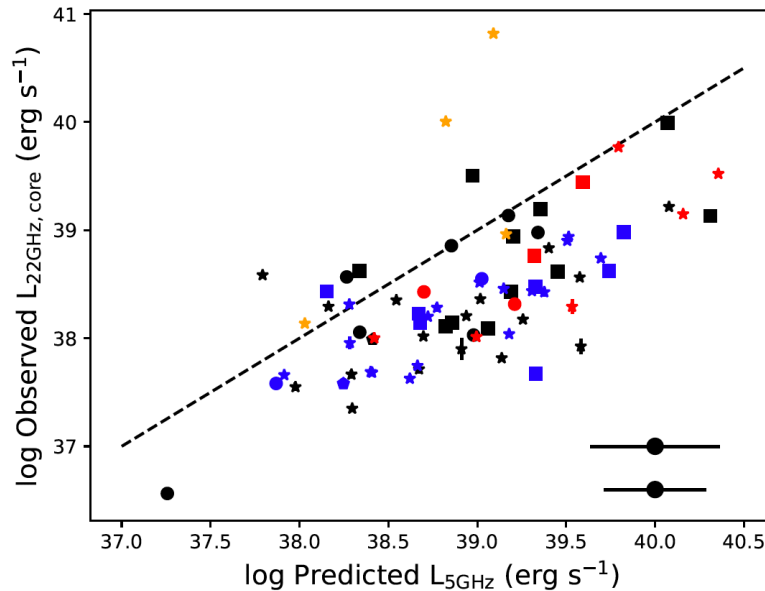
The core L_R / L_X is consistent with $\sim 10^{-5}$, the same as coronally active stars.



The “Fundamental Plane of Black Hole Activity”

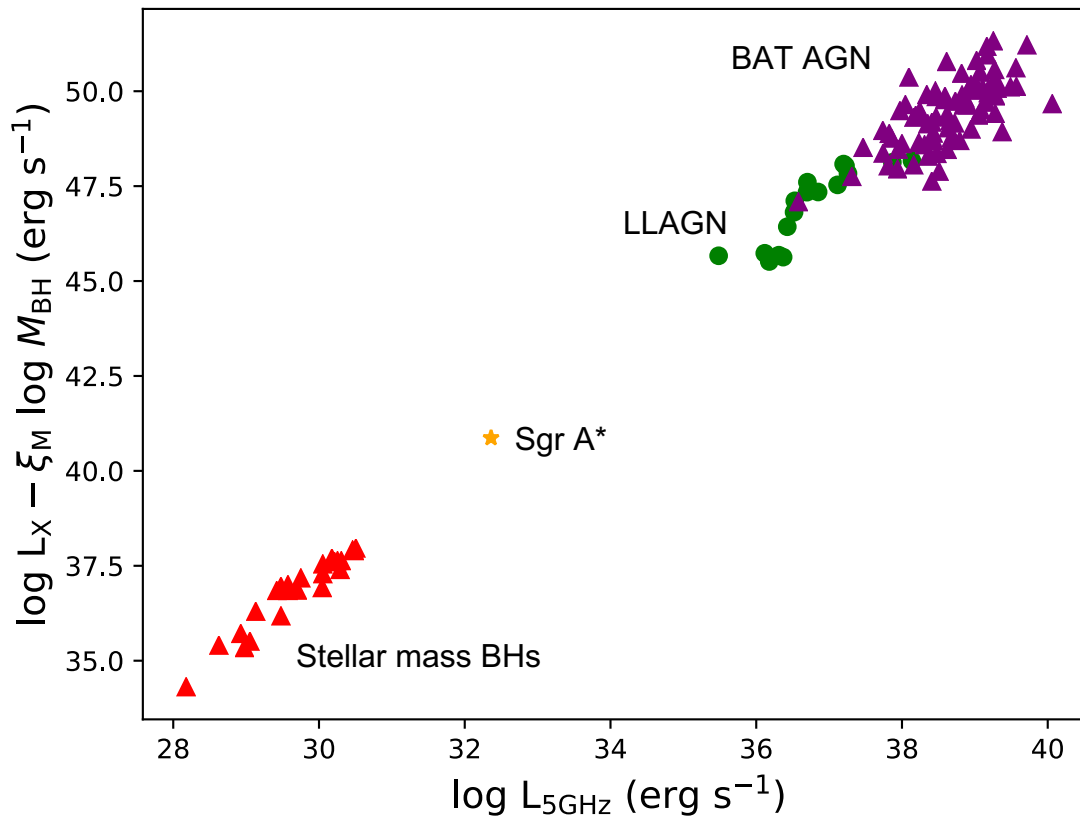
Merloni et al. 2003, Falcke et al. 2004

$$\log L_R = \xi_{RX} \log L_X + \xi_{RM} \log M + K$$



The BAT AGN sample lies below the plane, consistent with the smaller Smith+2016 sample and with Wong+ 2016.

But if you zoom way out...



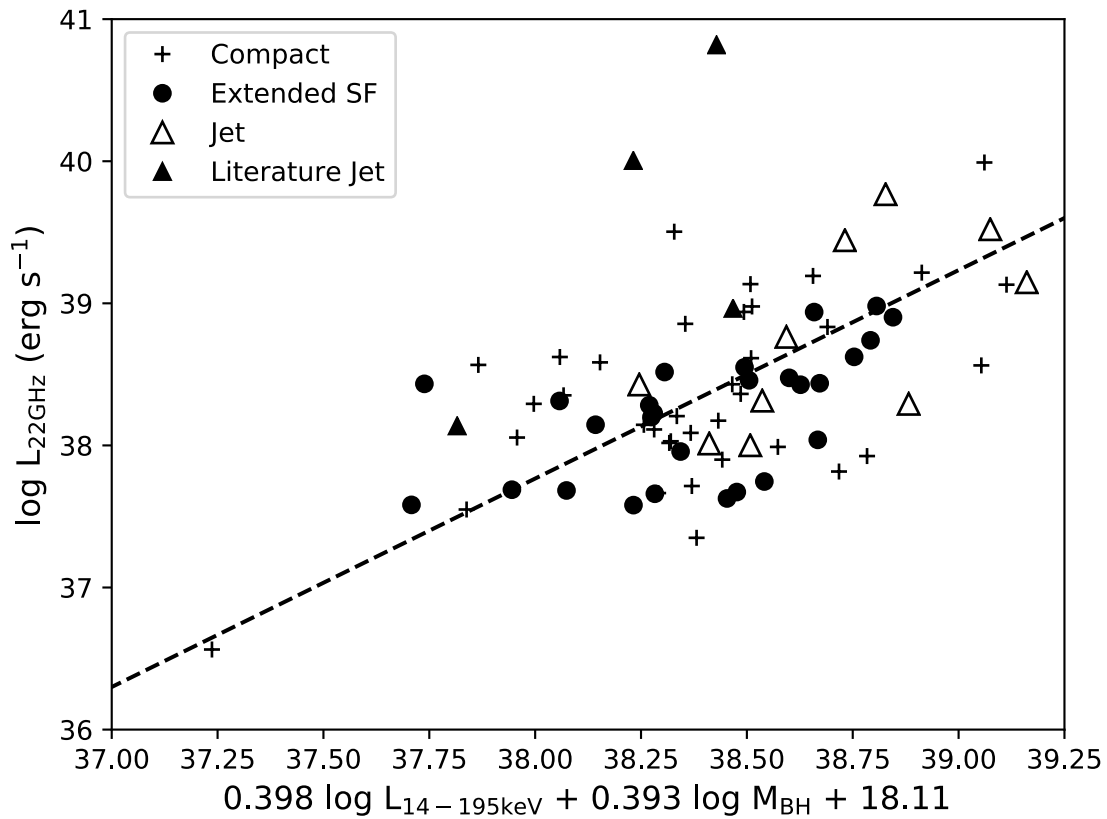
Plotkin et al. (2012) + core BAT AGN data


● A 22 GHz, UHX Fundamental Plane

○ ◦ May probe the true radio core better than lower frequencies:

- Synchrotron self-absorption region size scales as $R_{\text{pc}} \sim \nu^{-7/4}$ (Behar 2018)
- Ultra-hard X-rays are guaranteed to come from the AGN itself

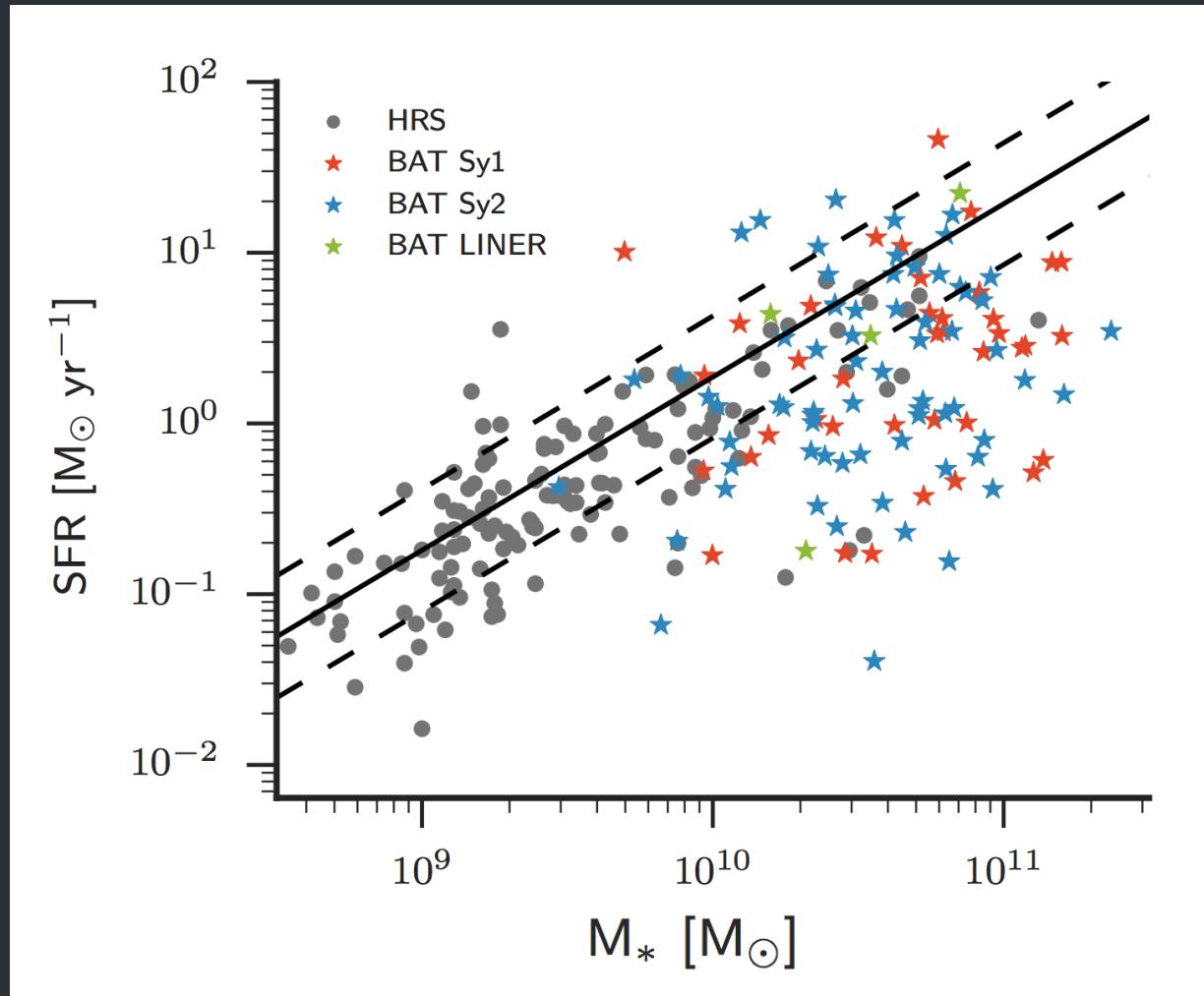
A 22 GHz, UHX Fundamental Plane



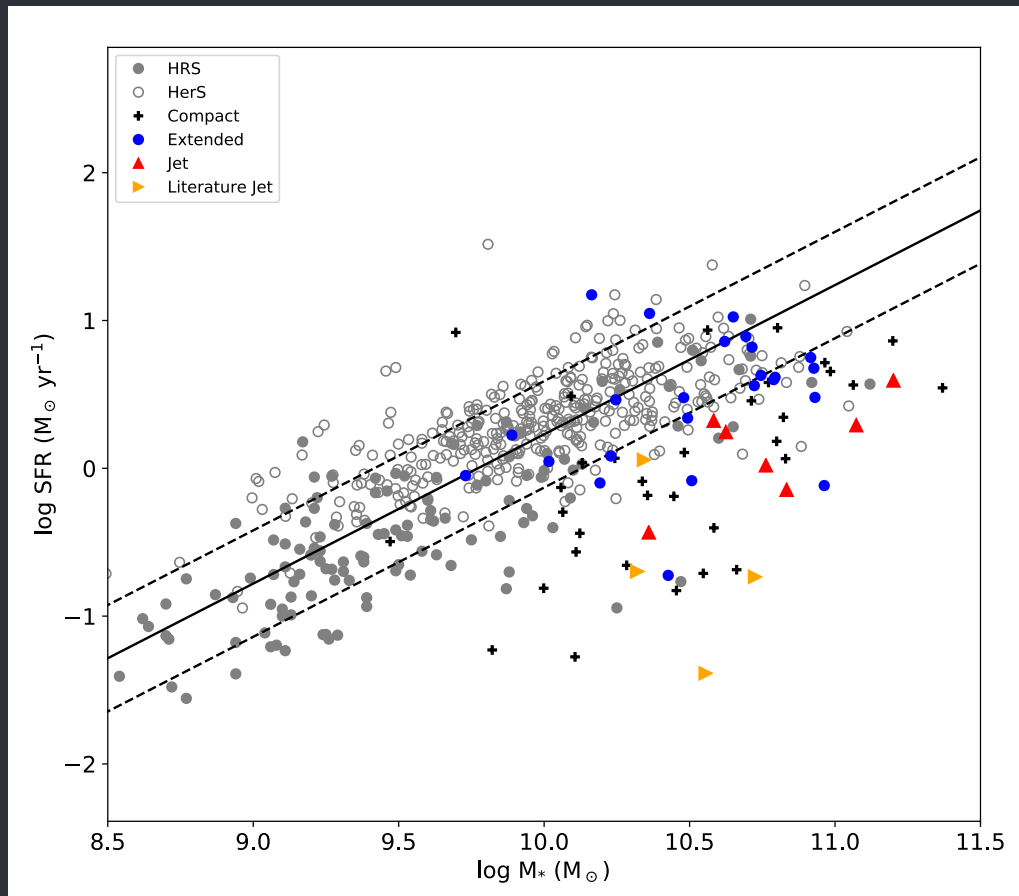
A vertical white line on the left side of the slide, with three small white circles positioned at regular intervals along its length.

**Back to star formation for a
moment...**

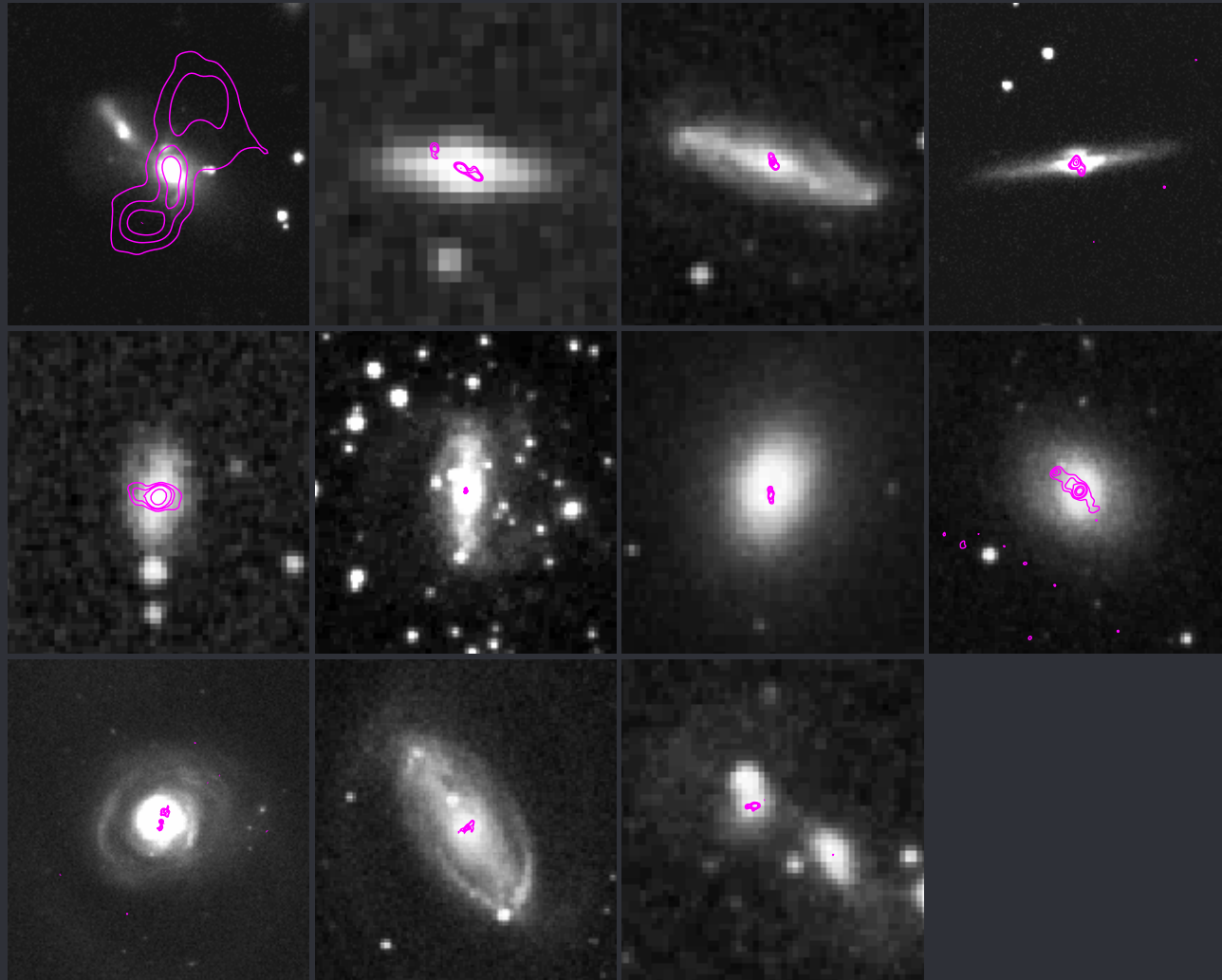
● Global star formation rate is suppressed in the BAT AGN.



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Status: Core paper in final internal review phase, submission in February.

Letter on jet / feedback result in drafting phase

The remainder of BAT AGN with *Herschel* imaging approved by JVLA with A-priority.

Phase I: 65 objects, completed December 2018

Phase II: 63 objects, propose in September

Result: A 22 GHz 1'' resolution atlas of the BAT AGN



Questions?